Session Program

Time	Speaker	Talk		
14:30	Hinder, Ian	A Review of Recent Developments in Numerical and Analytical Relativity		
15:00	Nagar, Alessandro	Gravitational waves from coalescing compact binaries: interfacing analytical and numerical techniques		
15:30	Le Tiec, Alexandre	The overlap of numerical relativity, perturbation theory and post-Newtonian theory in the binary black hole problem		
Coffee Break				
16:30	Calderón Bustillo, Juan	Accuracy of Complete Hybrid PN/NR Descriptions Of The Gravitational Radiation From Non-Precessing Compact Binaries		
17:00	Vano-Vinuales, Alex	Free Hyperboloidal Evolution In Spherical Symmetry		
17:30	Puerrer, Michael	Accelerating Parameter Estimation of Gravitational Waves from Black Hole Binaries with Reduced Order Quadratures		
18:00	Balmelli, Simone	The description of next-to-leading order spin-spin effects in an Effective-One-Body Hamiltonian		
Author		Poster (FISICA Marconi, all week)		
Author				
Celestino, Juliana		Nonlinear evolution of cylindrical gravitational waves		
llseven, Ekin		Lattice Boltzmann Model for Numerical Relativity		



Nonlinear evolution of cylindrical gravitational waves

- Juliana Celestino, Henrique P. de Oliveira, Eduardo L. Rodrigues
- There is currently a vivid expectation of a direct detection of gravitational ٠ waves from one or more detectors built specially to this aim. Meanwhile, theoretical studies concerning the wave-forms generated by sources of gravitational radiation and their efficiency are of great interest and constitute a valid effort to understand the role played by the nonlinear character of Einstein equations. The use of numerical techniques is unavoidable if one wants to accomplish any success in exploring at theoretical level the fundamental aspects of gravitational wave emission. In spite of their unphysical character, cylindrical waves constitute an interesting nonlinear problem in which both polarization modes are present. We have constructed an efficient numerical code to integrate the field equations based on a combination of Galerkin and pseudospectral methods. As usual the code tests were performed quite satisfactorily, and we have focused on the nonlinear evolution.

A Review of Recent Developments in Numerical and Analytical Relativity

Ian Hinder



Max Planck Institute for Gravitational Physics, (Albert Einstein Institute), Potsdam, Germany 16th July, 2015 14th Marcel Grossmann Meeting, Rome

Overview

- The Numerical Relativity field
- Analytic inspiral-merger-ringdown waveform families
- Edges of NR parameter space
- A selection of recent results

The Numerical Relativity field

- **118 papers** related to Numerical Relativity in the past year
- NR waveform catalogues:
 - NINJA 1, 23 waveforms, 2009
 - NINJA 2, 60 hybrid waveforms, 2012
 - NRAR, 25 waveforms, 2013
 - SXS: Originally 174 waveforms. Now 201. 2013 onwards.

Latest "Analytic" waveform models

- Effective-One-Body variants:
 - Precessing SEOBNRv3 (Taracchini et al. 2014) and Damour and Nagar, 2014 (recalibrated June 2015)
 - System of ODEs solved via numerical integration
 - Computational cost can be decreased via use of reduced-order modelling (see Michael Pürrer's talk today)
- Precessing **PhenomP** Hannam et al. 2014:
 - "Simple" closed-form frequency-domain expression for h; singlespin and stationary-phase approximations

Latest "Analytic" waveform models

• **PhenomD** (in preparation, **PRELIMINARY**):

- 19 calibration points in q, S space including q=18, χ₁=-0.8
- Mismatch < 1% over calibration range for 47 waveforms



Edges of NR parameter space

- Number of cycles / lowest frequency
- High spin
- High mass ratio

Longest NR waveform

- Szilagyi et al., 2015 (Numerical relativity reaching into post-Newtonian territory: a compact-object binary simulation spanning **350 gravitational-wave cycles**)
- SpEC code
- Mass ratio q=7
- Separation D = 27 M
- NR waveform covers aLIGO band for M₁+M₂ >
 45.5 M_☉



 $(t - t_{\text{peak}}) / 1000M$

- Previous longest: 60 cycles
- Pan et al., 2013: with NR errors at the time, 30 and 50 cycle EOB waveforms were indistinguishable to aLIGO



- Started 4000 GW cycles before merger
- Evolved for 3 cycles (no merger) at D = 100 M

TABLE IV: Energy and angular momentum radiated per orbit (initial and second).

D	$\delta m/M_{ m num}$	$\delta m/M_{ m PN}$
20M	$(5.68 \pm 0.02) \times 10^{-5}$	$5.43 - 5.62 \times 10^{-5}$
50M	$(2.4 \pm 0.1) \times 10^{-6}$	$2.52 - 2.53 \times 10^{-6}$
100M	$(2.3 \pm 0.4) \times 10^{-7}$	$2.36 - 2.36 \times 10^{-7}$
D	$\delta J/M_{ m num}^2$	$\delta J/M_{ m PN}^2$
20M	$(5.39 \pm 0.01) \times 10^{-3}$	$5.20 - 5.30 \times 10^{-3}$
50M	$(8.9 \pm 0.1) \times 10^{-4}$	$9.16 - 9.18 \times 10^{-4}$
100M	$(2.4 \pm 0.4) \times 10^{-4}$	$2.39 - 2.39 \times 10^{-4}$

Highest spins



- Scheel at al. 2014 (Improved methods for simulating nearly extremal binary black holes)
- Superposed Kerr-Schild initial data to exceed the Bowen-York limit
- SpEC code with improved numerical methods and code optimisation
 - Simulations:
 - q=1, aligned χ₁=χ₂=0.994, 50 cycles
 (highest spin ever simulated)
 - q=1, aligned **χ**₁=**χ**₂=**0.99**, 50 cycles
 - q=1.5, χ₁=0.99 (aligned), χ₂=0.2 misaligned, 46 cycles
 - NR phase error smaller than difference with PN and EOB approximants for χ =0.994



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Highest mass ratio

- **q = 100** (1 orbit), Nakano et al., 2011 (Intermediate-mass-ratio black hole binaries II: Modeling Trajectories and Gravitational Waveforms)
- **q = 18**, 7 orbits, non-spinning (Varma et al., 2014)
- [Preliminary] q = 18, 8 orbits, $\chi_1 = -0.8$



Recent Results

Energy and Angular Momentum Comparison

- Nagar et al., 2015 (Energetics and phasing of nonprecessing spinning coalescing black hole binaries)
- Radiated E and J computed from NR simulations using SpEC and Llama codes and compared with EOB

See talk by Alessandro Nagar (next)



Precession Dynamics

- Ossokine et al., Jan 2015
- 37 precessing simulations (from SXS catalogue)
- Mass ratios up to 8, up to 2
 precession cycles (>~ 60 wave cycles)
- Orbital angular momentum and spins agree between NR and PN to within 1 degree during inspiral and within 5 degrees at merger



Sub-dominant modes



Spin "flip-flop"

- Lousto and Healy,
 2015 (Flip-Flopping Binary Black Holes)
- Lousto and Healy, 2015 (Spin flips in generic black hole binaries)
- LazEv code
- Reproduction of PN spin dynamics using NR



FIG. 2. The angle between the spin of the secondary black hole \vec{S}_1 with respect to the orbital angular momentum \vec{L} (left) and with respect to the fixed z-axis (right). For comparison we also plot the 3.5PN prediction.

Cosmology: Black Hole Lattices

 Korzyński, IH, Bentivegna 2015 (On the vacuum Einstein equations along curves with a discrete local rotational and reflection symmetry)



Claim in literature: Einstein equations **reduce to ODEs** for metric and electric Weyl on edge:



Lattice cell edges after stereographic projection to R³

but full 3D numerical relativity evolution disagrees (NR error bars negligible in plot)

Cosmology: Black Hole Lattices

 Careful re-examination of ODE reduction argument reveals incorrect symmetry assumption. Derived missing term in ODE:

$$\frac{\ddot{a}_{\parallel}}{a_{\parallel}} = \frac{2}{3}E_{+}, \qquad \frac{\ddot{a}_{\perp}}{a_{\perp}} = -\frac{1}{3}E_{+}, \qquad \dot{E}_{+} = -3\frac{\dot{a}_{\perp}}{a_{\perp}}E_{+} - \frac{3}{2}U_{11}, \qquad U_{11} = (K_{ij;k}{}^{k} - K_{ik;j}{}^{k}) \mathbf{e}_{1}^{i} \mathbf{e}_{1}^{j}$$

 NR shows that this perfectly accounts for the anomaly

$$-\frac{2}{3}\mathcal{A} = 1.00000(3)\,U_{11}$$

- A tedious Mathematica calculation eventually proved analytically that U₁₁ has a non-zero **3rd time derivative** at t=0. Discrepancy grows as t^{6.}
- Verified value with NR to 4 digits of precision

$$\frac{\partial^3 U_{11}}{\partial t^3}\Big|_{t=0,\phi=0} = \begin{cases} 4.3015(4) \times 10^{-12} \mathcal{M}^{-6} & \text{Numerical} \\ 4.30113 \times 10^{-12} \mathcal{M}^{-6} & \text{Analytic}. \end{cases}$$



(c) $\Delta \equiv 1 - (-2/3\mathcal{A})/U_{11}$ computed from NR at several resolutions.

Summary

- NR BBH simulations up to q=18, or 350 cycles, or S/m²=0.994. Pick any one!
- Recent results on:
 - Energy and Angular Momentum Comparison
 - Precession Dynamics
 - Sub-dominant modes
 - Spin "flip-flop"
 - Cosmology: Black hole lattices